



# **Micrometeoroid and Orbital Debris Impact Inspection of the HST WFPC2 Radiator**

**J.-C. Liou and the HST WFPC2 MMOD  
Inspection Team**



## Objectives

- **Conduct micrometeoroid and orbital debris (MMOD) impact inspection of the Hubble Space Telescope (HST) Wide Field Planetary Camera 2 (WFPC2) radiator**
  - Document the physical characteristics of large ( $\geq 300 \mu\text{m}$ ) impact features
    - Locations and distribution
    - Crater shape, size, depth, volume
  - Extract residues from craters for compositional analysis
  - Perform hypervelocity impact tests for data interpretation
- **Use the data to validate or improve the near-Earth MMOD environment definition**
  - Flux
  - Orbital debris versus meteoroids
  - Time history (limited)



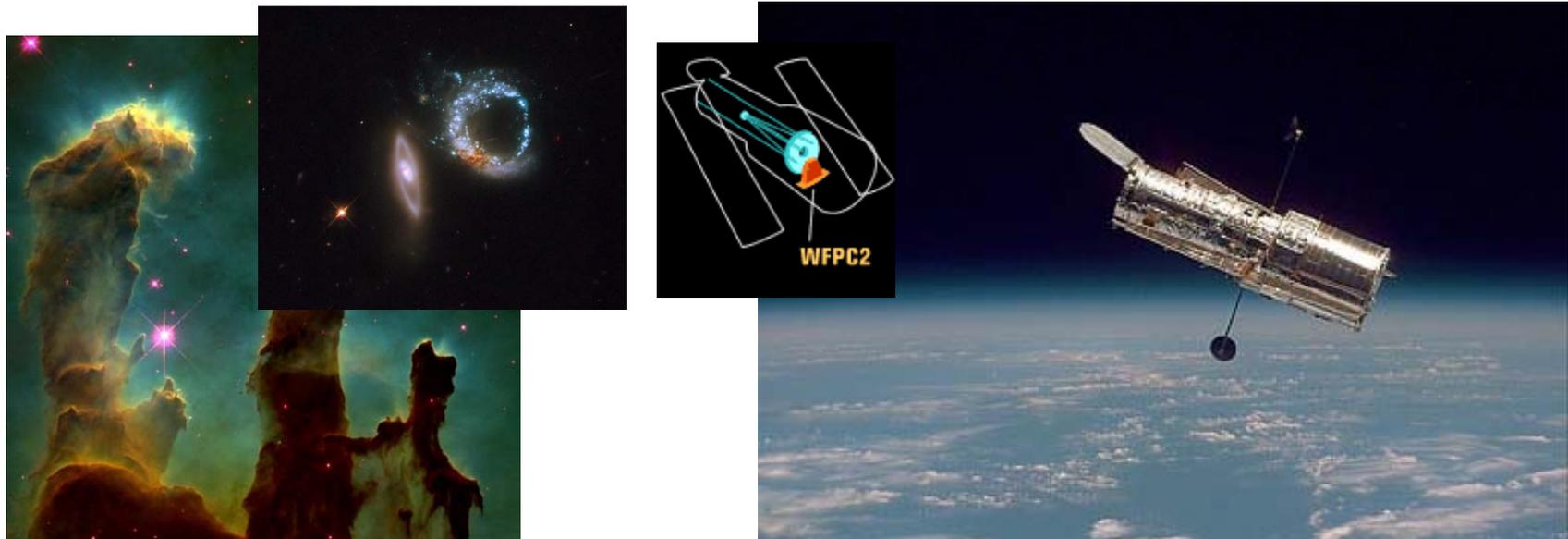
## Team

- **NASA Orbital Debris Program Office (lead)**
  - J.-C. Liou, P. Anz-Meador, J. Opiela, E. Stansbery, et al.
- **NASA Curation Office**
  - K. McNamara
- **NASA Hypervelocity Impact Technology Facility**
  - E. Christiansen, T. Hedman, J. Hyde, D. Ross, *et al.*
- **NASA Meteoroid Environment Office**
  - D. Moser, D. Edwards
- **NASA HST Program**
  - B. Reed



# History

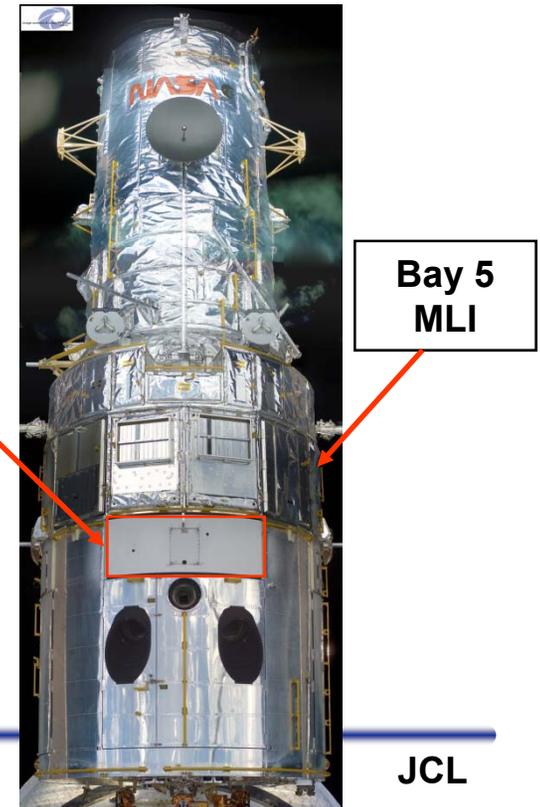
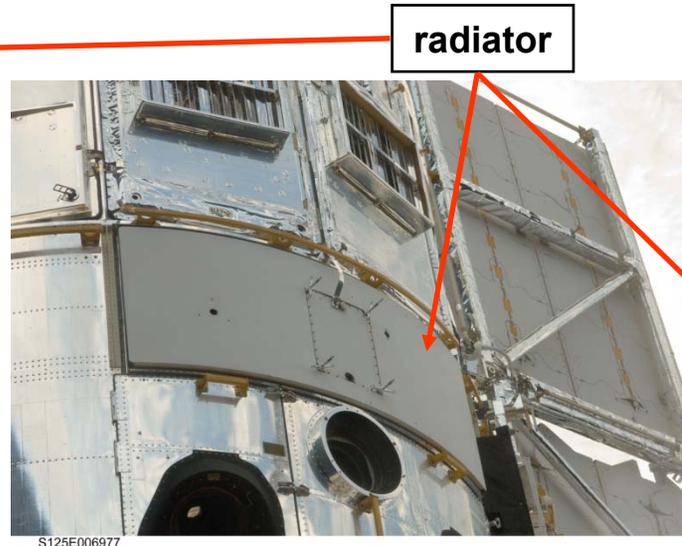
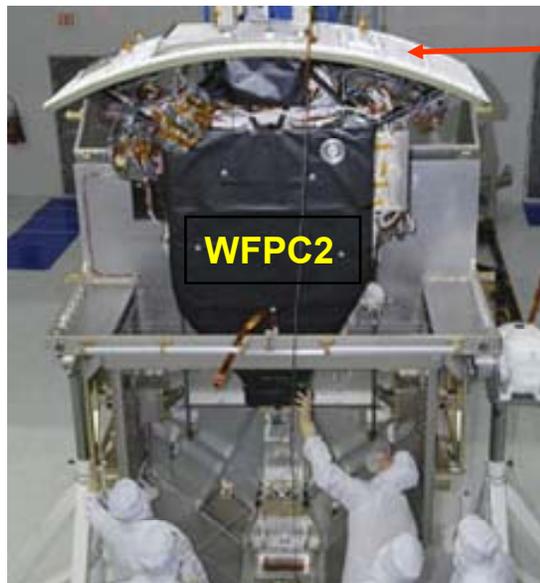
- **HST was launched from Discovery on 24 April 1990**
- **WFPC1 was replaced by WFPC2 during STS-61 HST Servicing Mission 1 (SM1) in December 1993**
  - **WFPC2 is the "workhorse" instrument behind nearly all of the most famous Hubble pictures**
- **WFPC2 was replaced by WFC3 during STS-125 HST SM4 in May 2009**





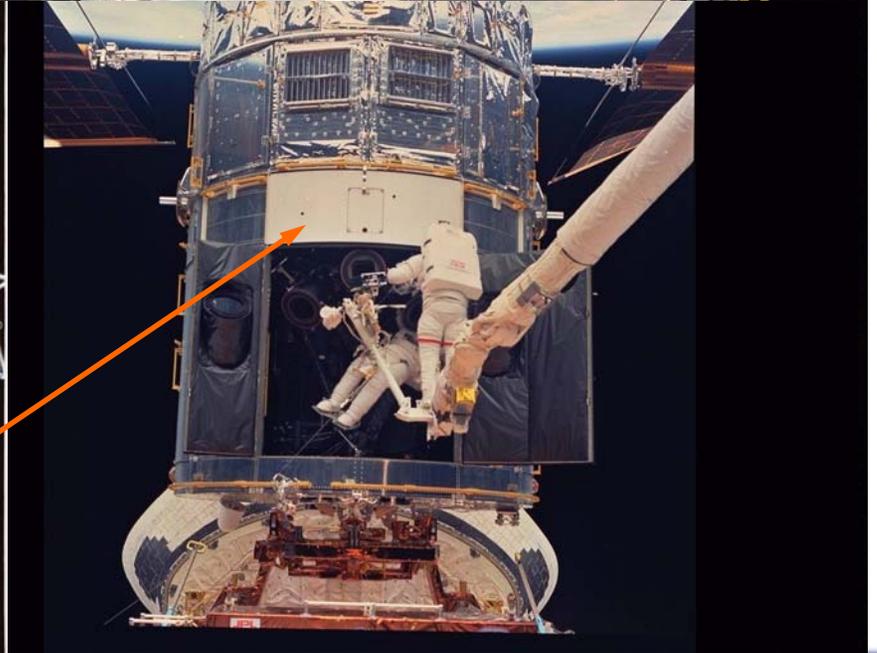
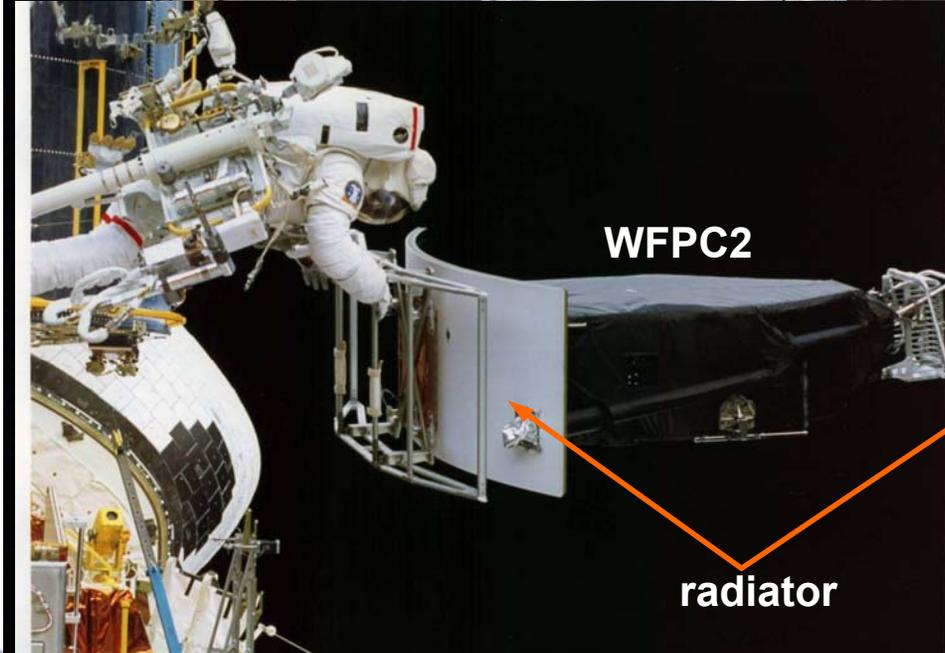
# WFPC2 Radiator Background

- The WFPC2 radiator was in space for 15.5 years (3.6 years for WFPC1 radiator)
- Dimensions of the radiator: 0.8 m × 2.2 m
- Outer layer: an aluminum plate (4.06 mm thick) coated with 4~8 mils Zinc Orthotitanate (ZOT, a ceramic thermal control paint)



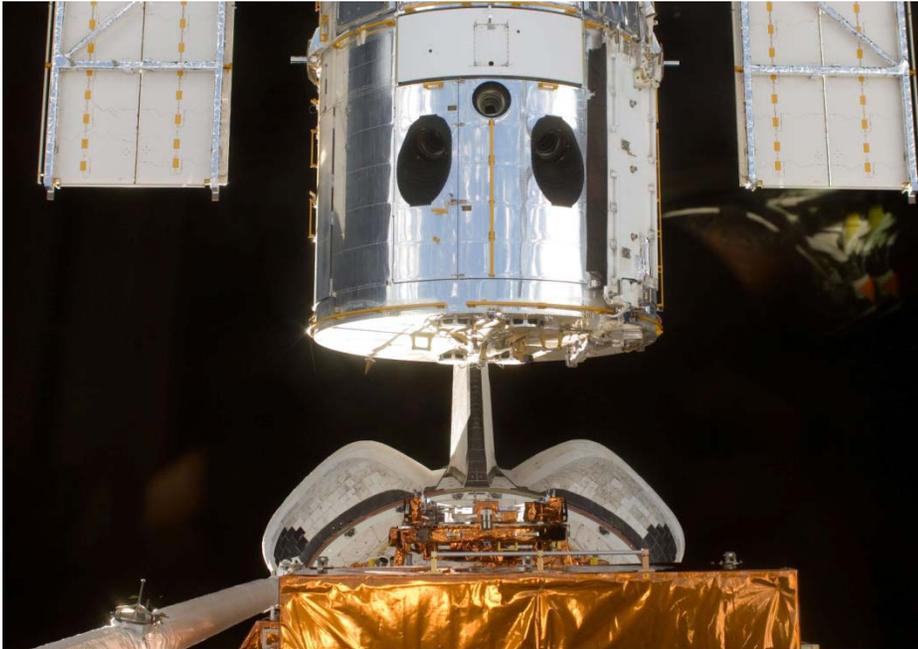


# HST SM1 (STS-61, 1993)





# HST SM4 (STS-125, May 2009)



S125E006992



S125E007068

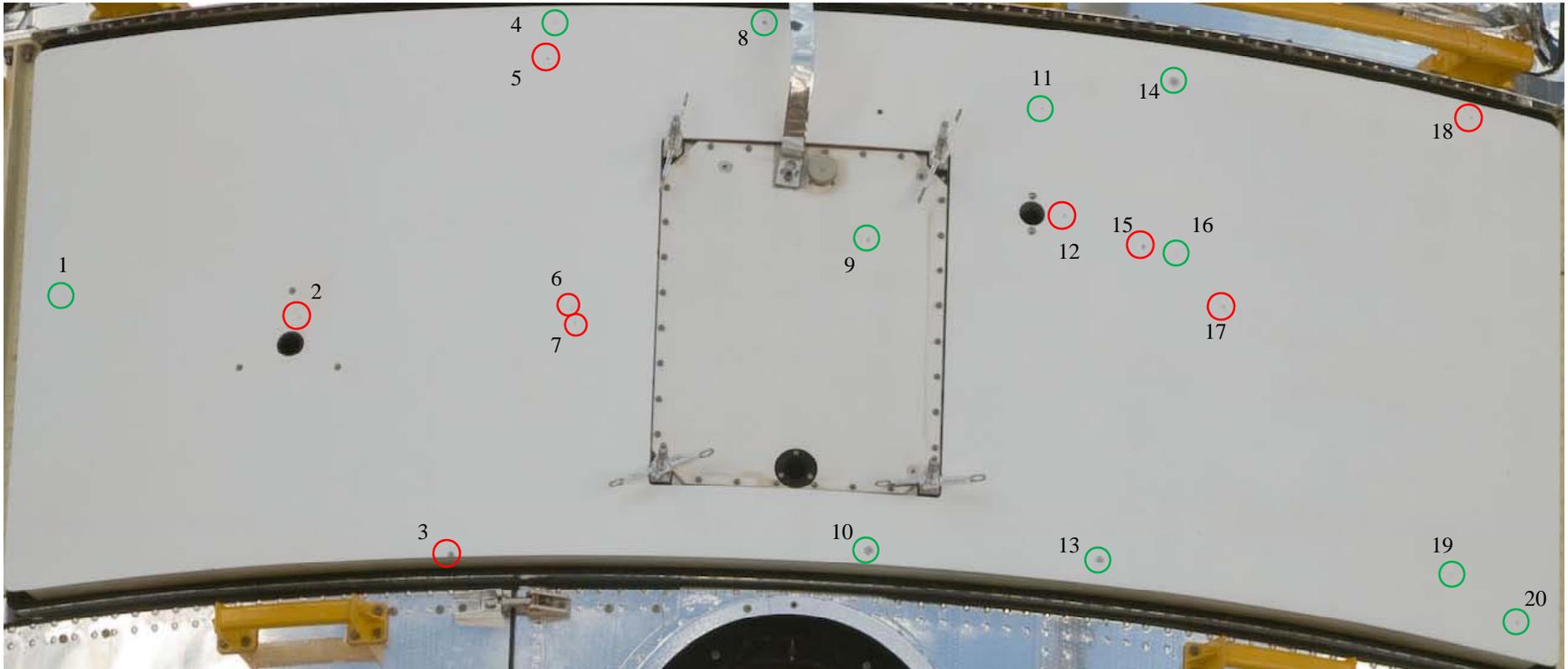


S125E006996



S125E007168

# Visible MMOD Impacts from the On-orbit Imagery Survey



S125e006995.jpg (edited)

- Red circles: Impacts identified from SM3B images (2002)
- Green circles: Impacts identified from SM4 images (2009)

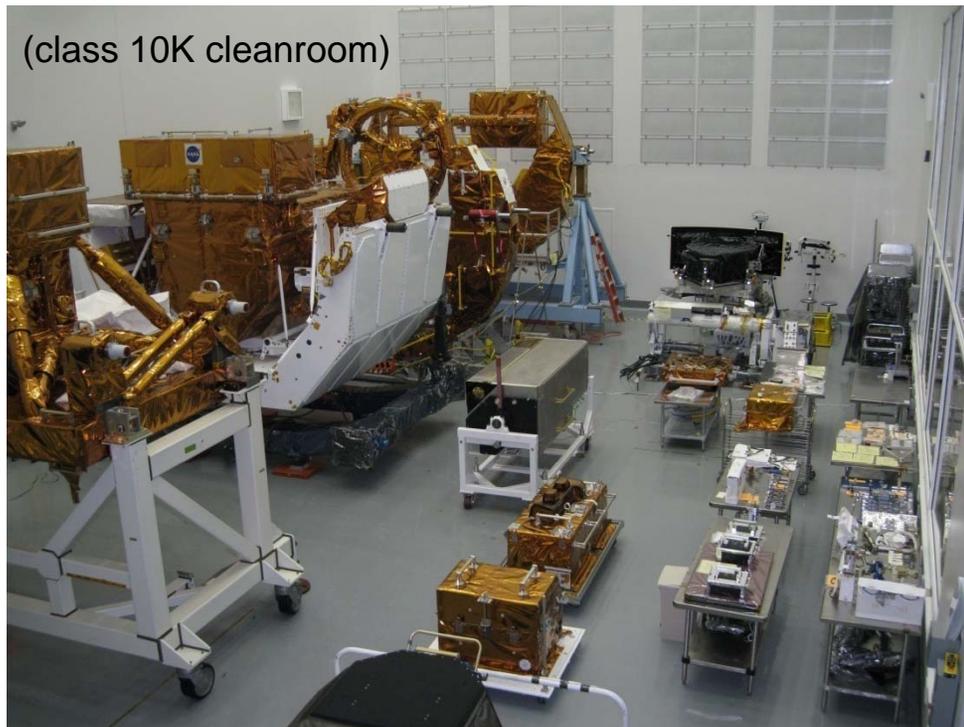
# Post-Flight Deintegration at KSC





# MMOD Inspection at NASA Goddard

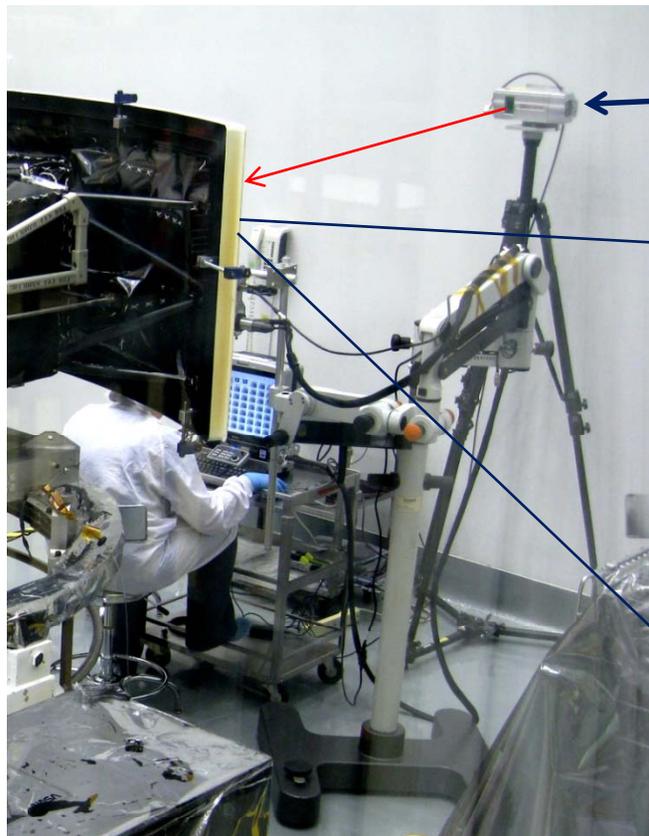
- **WFPC2 was shipped to Goddard in late June 2009**
- **MMOD inspection: 6-17 July, 24 August – 4 September, and 14-25 September 2009**
  - **Major inspection instruments: LAP CAD-Pro laser template projector and VHX-600 digital microscope**





# Inspection Instruments (2/3)

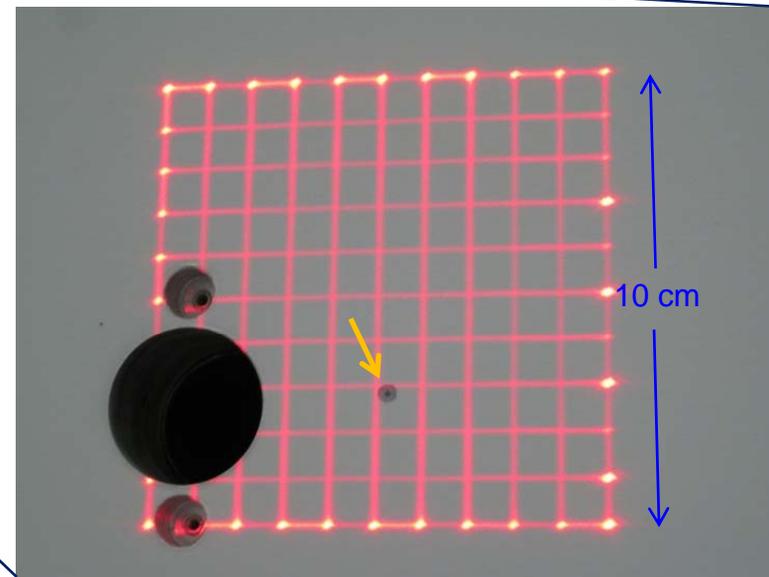
- **LAP CAD-Pro laser template projector**
  - Provide crater coordinates for VHX-600 operator



Cleanroom setup



Laser projector head

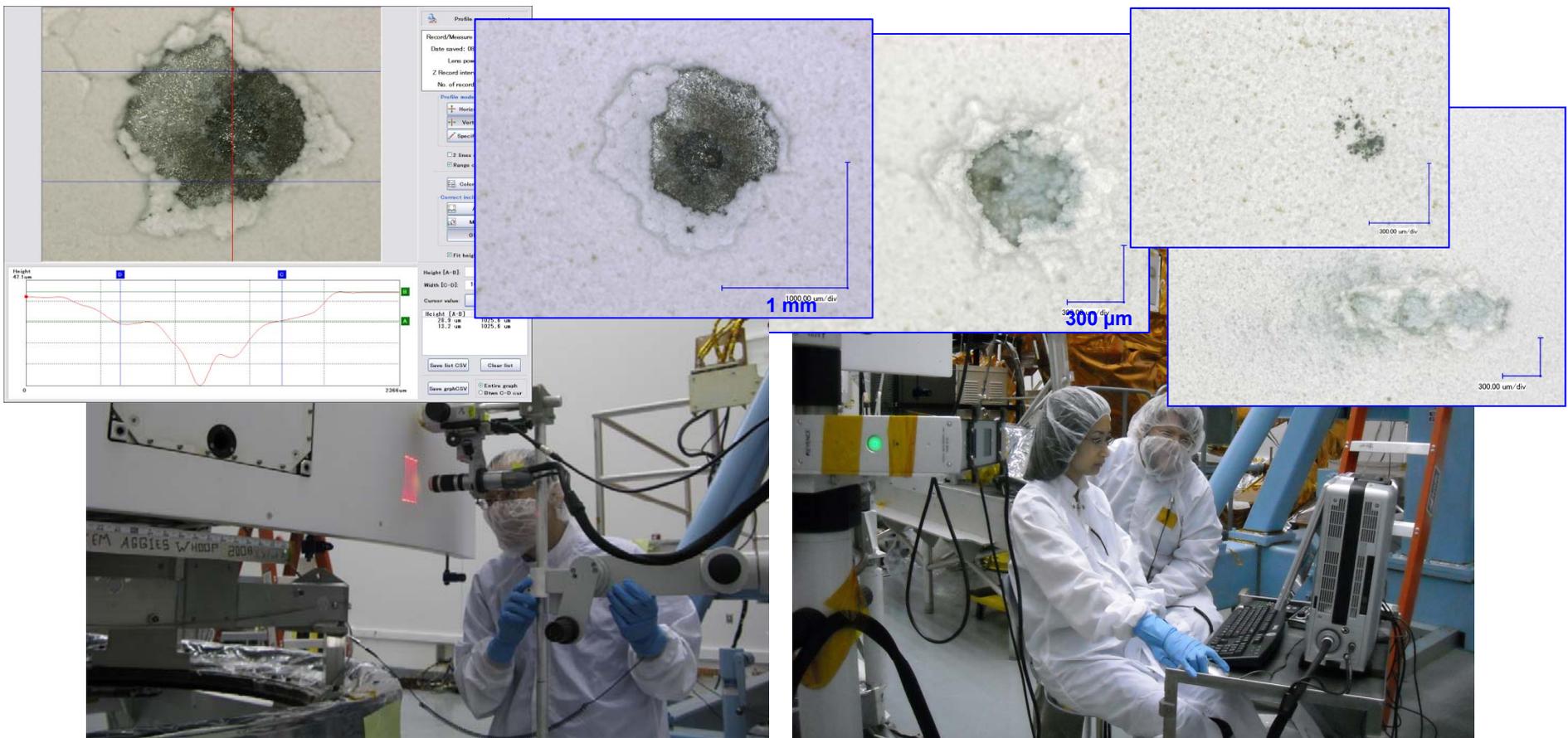


Square grid projected on a curved surface



# Inspection Instruments (3/3)

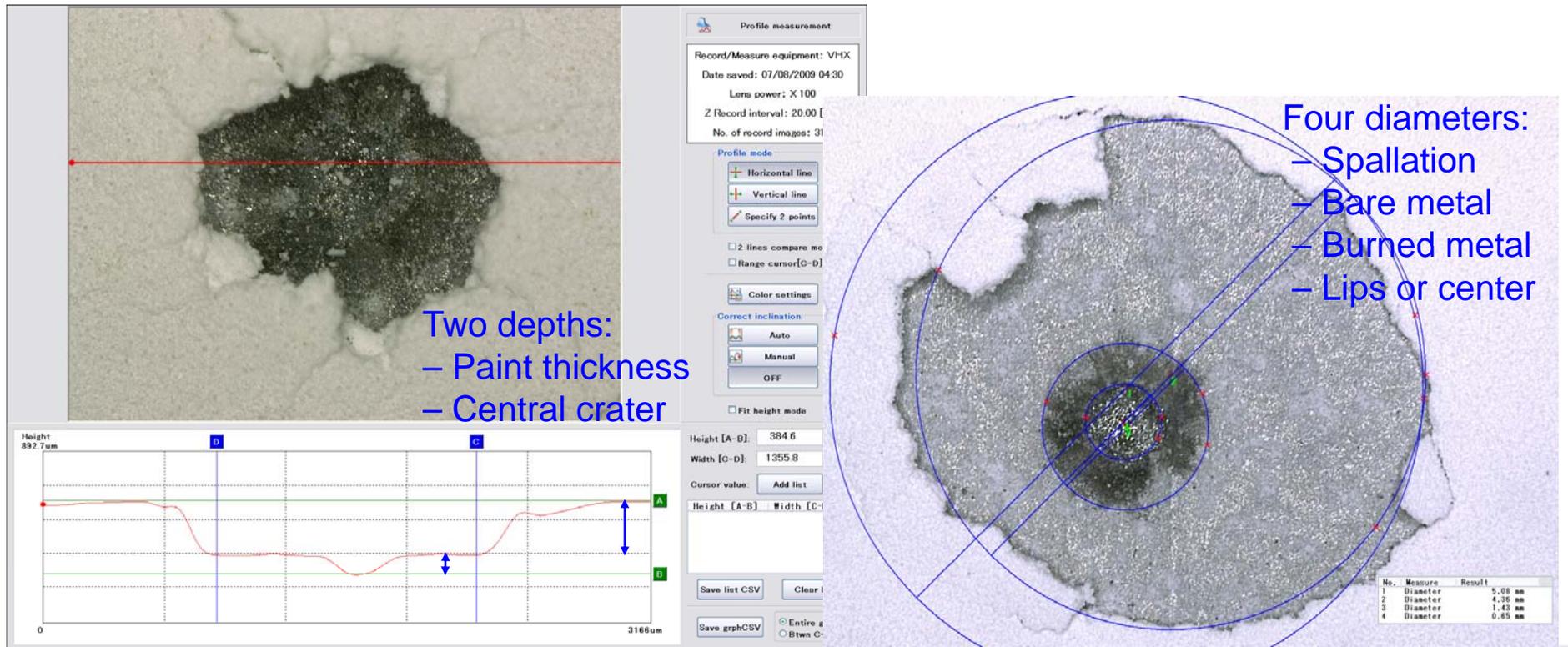
- **Keyence VHX-600 digital microscope (up to 5000x optical magnification, 2-D and 3-D imagery)**
  - Record each impact feature's shape, size, depth, and volume





# Inspection Results and Data Processing

- Documented 685 impact craters ( $\geq 300 \mu\text{m}$ ) and numerous non-impact features
- No through-hole
- The largest one: 1.6 mm crater plus 1.4 cm spall zone

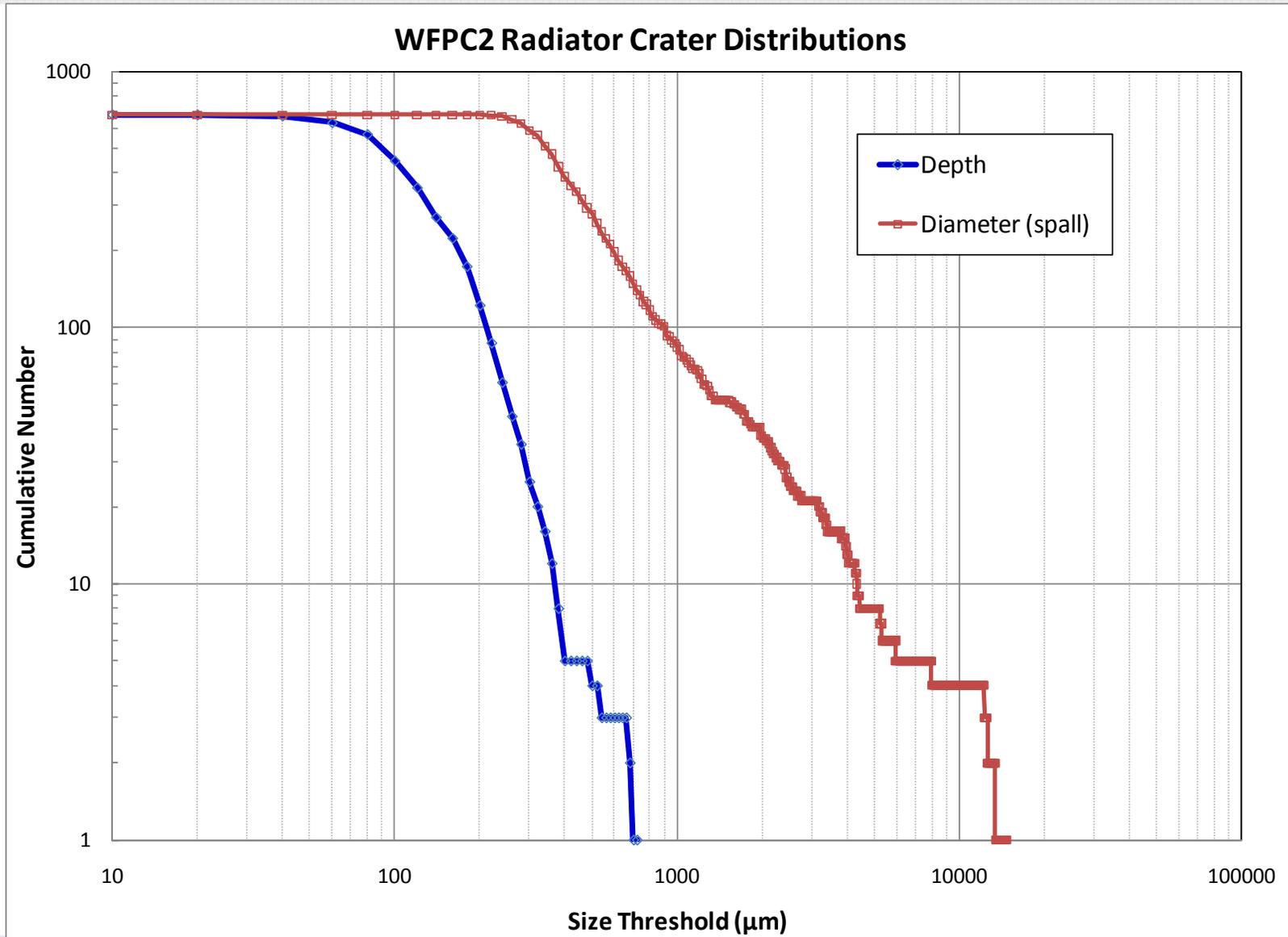


Two depths:  
 - Paint thickness  
 - Central crater

Four diameters:  
 - Spallation  
 - Bare metal  
 - Burned metal  
 - Lips or center



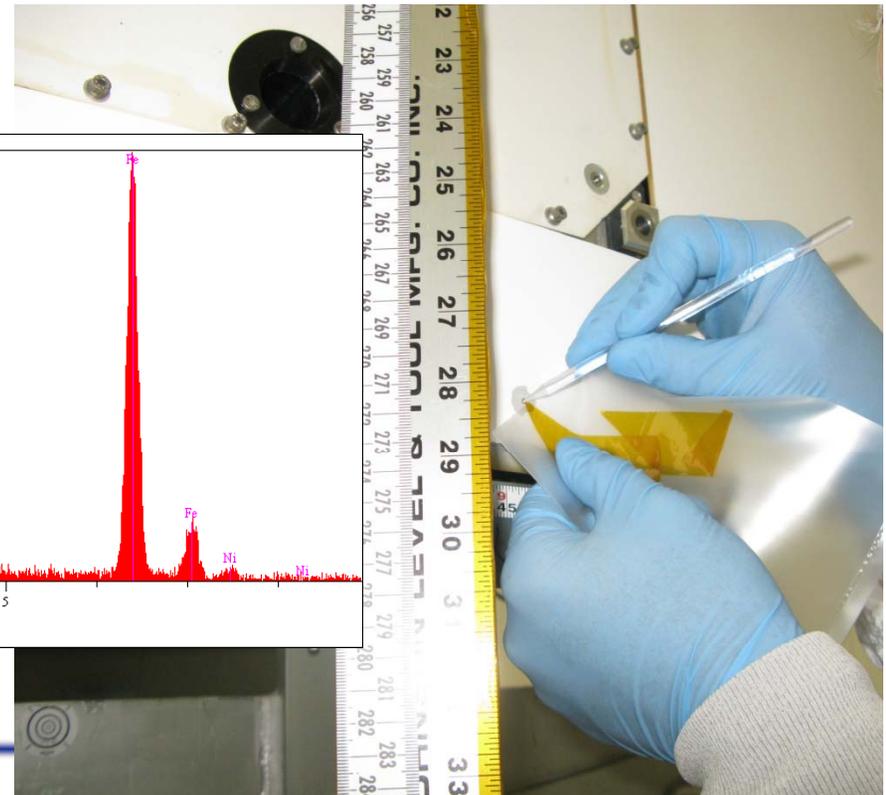
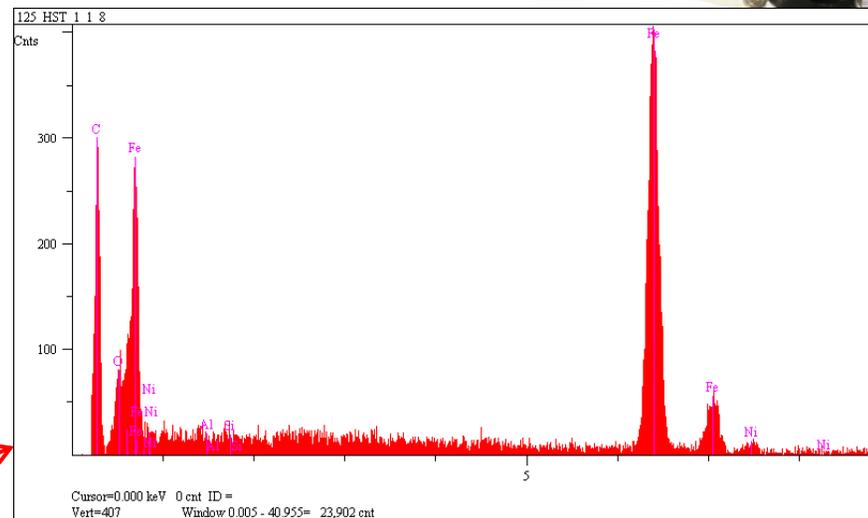
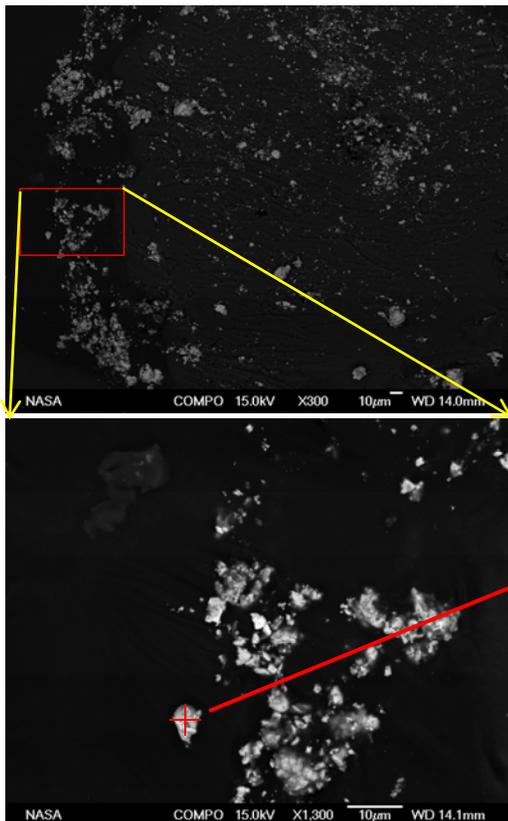
# Radiator Crater Data





# Sample Collection

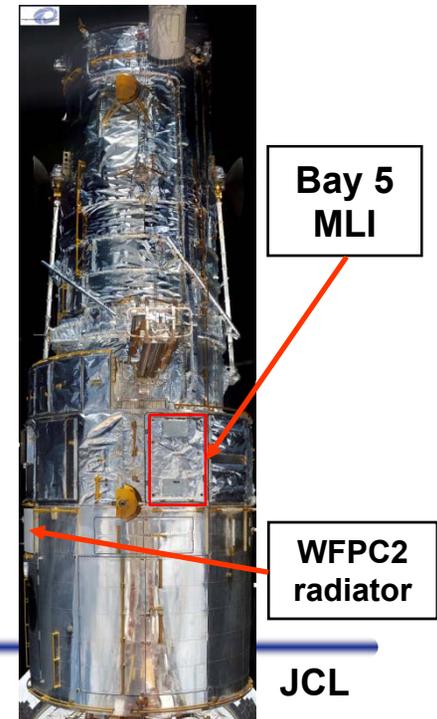
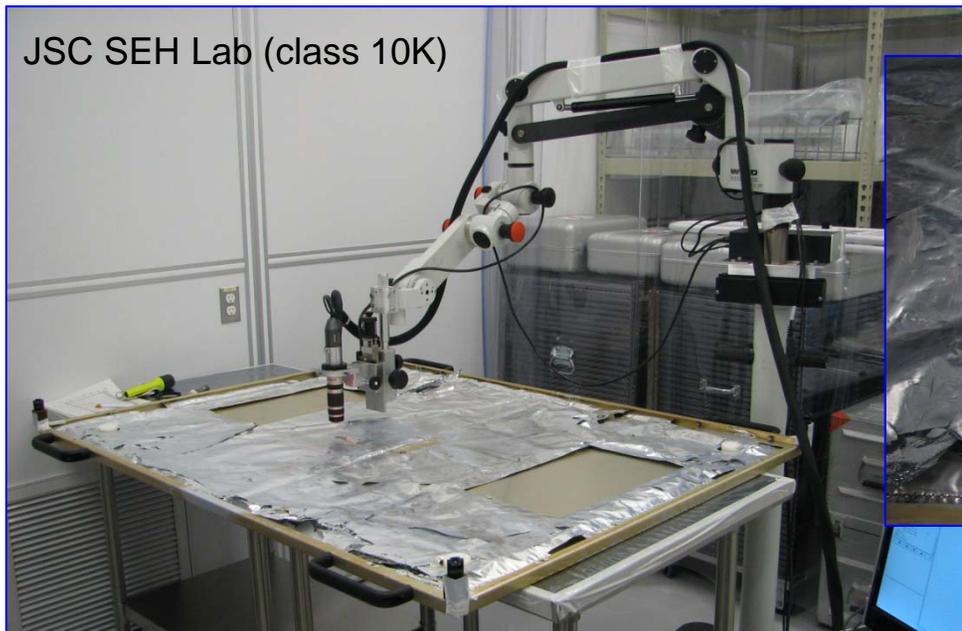
- Collected tape pull samples from 50 features in July
- Conducted preliminary scanning electron microscope (SEM) analysis on 36 samples
  - Identified non-radiator materials (Fe-Ni, Mg-silicate, steel, etc.) from 8 of them





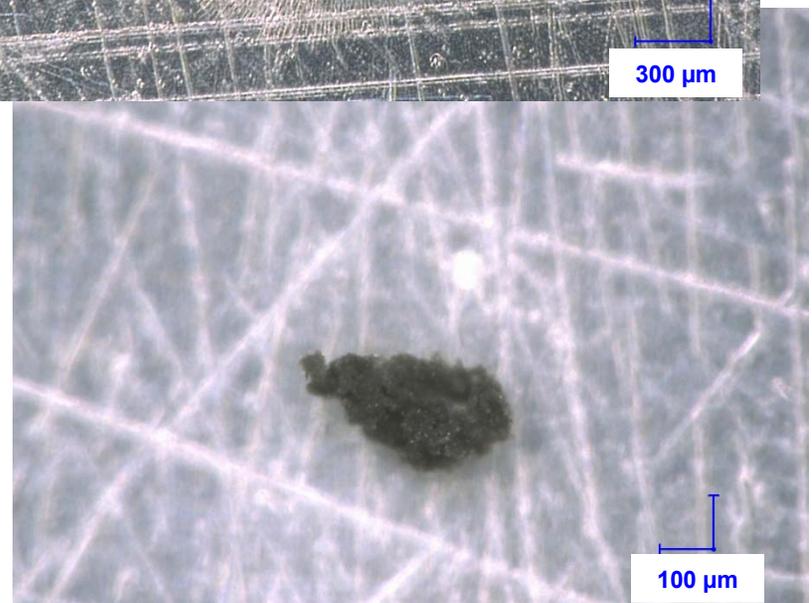
# Bay 5 Multi-Layer Insulation (MLI) Inspection

- **The HST Bay 5 MLI was shipped to JSC for 5-week MMOD inspection on 2 Feb 2010**
  - Bay 5 MLI was in space between 1990 and 2009
  - Dimensions of MLI: 1.1 m × 1.5 m (with several cut-out areas)
  - Consists of 17 layers
    - The outermost layer: 127 μm thick FEP (fluorinated ethylene-propylene) Teflon with vapor deposited Al (VDA) coated on the backside
  - Document impact features as small as ~100 μm





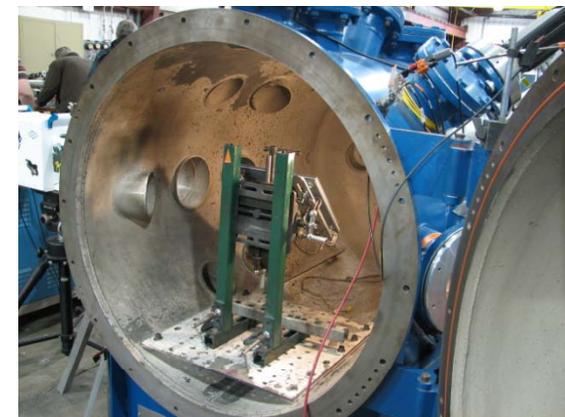
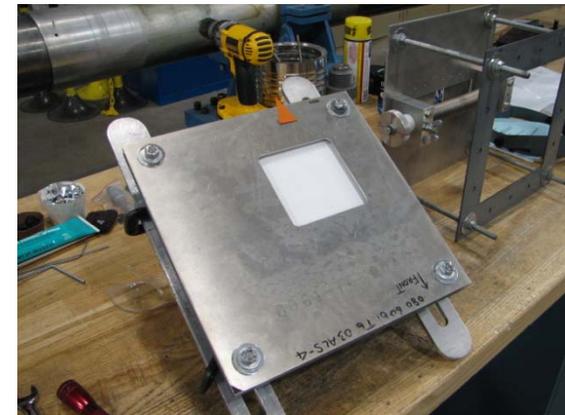
# Sample MLI Inspection Data





# Hypervelocity Impact Tests

- **60 shots scheduled for projectiles of different sizes, materials, impact speed, and impact angle**
  - Started on 17 Feb 2010



.17 caliber 2-stage light gas gun at JSC White Sands Test Facility



## Forward Plan

- **Proceed with core sample proposal/plan**
- **Complete hypervelocity impact tests**
  - Process and analyze crater data, as with WFPC2
  - Establish feature-to-projectile conversion
- **Conduct hydrocode simulations**
  - Supplement hypervelocity impact data
  - Extrapolate to impact speed above 8 km/sec
- **Process and analyze MLI data**
- **Model the orbital debris and micrometeoroid environment**
  - Separate orbital debris from micrometeoroids
  - Compare with previous solar array data
- **Document and publish the results**



## Value of the Radiator Impact Data

- **The NASA Orbital Debris Engineering Model (ORDEM)**
  - Describes the near-Earth OD environment
  - Is used by NASA (STS, ISS, *etc.*), DoD, and other national and international groups for satellite impact risk assessments and shielding designs
- **The database for the 100- $\mu$ m-to-1-mm particles in the new ORDEM to be released in 2010**
  - 211 impacts from 35 STS missions (1995-2002) at 350 to 400 km altitude
  - 17 impacts from 3 STS missions at 560 to 620 km altitude
- **Number of impact craters corresponding to particles in the same size regime from the WFPC2 radiator: 685**
  - Triple the total database
  - A 40x increase for data at high altitude